

CHM 5423 – Atmospheric Chemistry

Problem Set 1

Due date: Thursday, January 24th.

Do the following problems. Show your work.

1) The concentration of molecular oxygen (O₂) for Mars at surface level is 1300 ppm. The total pressure and temperature of the atmosphere at surface level is $p_{\text{total}} = 0.0060$ bar, $T = 210$ K. Give the amount of oxygen present at surface level in the following ways:

- $p(\text{O}_2)$ (partial pressure of oxygen, in units of bar, atm, torr, and μtorr .)
- $N(\text{O}_2)$ (number density of oxygen molecules, in units of molecules/cm³)
- $D(\text{O}_2)$ (density of oxygen, in units of $\mu\text{g/L}$)

2) The following question concerns the escape of molecules from the atmosphere of Mars.

a) Find v_{esc} and v_{rms} for the following molecules: H₂, H₂O, N₂, and Ar. Use $T = 500$ K (an approximate average temperature in the Martian exosphere) in your calculations.

b) Define $R_v = v_{\text{esc}}/v_{\text{rms}}$, the ratio of the escape velocity to the rms average speed of a molecule. It can be shown that for the case $R_v \gg 1$ that f , the fraction of molecules in a gas at equilibrium that have a speed great enough to leave the gravitational attraction of a planet is given by the expression

$$f \cong (1/6\pi R_v^2)^{1/2} \exp(-3R_v^2/2) \quad (2.1)$$

Find f for each of the above four molecules in the Martian exosphere.

c) The likelihood of molecules of a particular gas escaping from the gravitational attraction of a planet can be estimated based on the value for f . In general, the larger the value for f the more likely escape of the gas becomes. We may use the following approximation to estimate the likelihood that a particular gas has escaped the gravitational attraction of a planet during the 4.5 billion years that the solar system has existed.

If $f < 10^{-25}$, then essentially no escape will have occurred.

If $10^{-15} < f < 10^{-25}$, then some escape will have occurred

If $f > 10^{-15}$ then essentially all of the gas will have escaped

Based on your results above, explain why the current atmosphere of Mars contains large amounts of N₂ and Ar but only small amounts of H₂.

d) There is good reason to believe that the average temperature of the Martian exosphere once was much larger than the value currently observed. Explain the implications of this for the escape of gases from the atmosphere of Mars.

e) The above argument for gas escape ignores the possibility of photochemically assisted escape of gases. For water, for example, the following process can occur in the exosphere



Explain how the above process could lead to the removal of water from the atmosphere of Mars.

3) At what altitude (in km) is the pressure of the Earth's atmosphere equal to the surface pressure on Mars (0.0060 bar)? Use $H = 7.4$ km for the scale height for Earth.

4) One indirect piece of evidence indicating that free molecular oxygen was uncommon in the early atmosphere of the Earth is the presence of pyrites (metal sulfides) in mineral samples that have been dated several billion years old. Such compounds are thermodynamically unstable in the presence of oxygen. More recent mineral samples have iron oxides present, but usually do not contain iron pyrite.

a) Using the information below, find the equilibrium constant for the process



b) Based on your answer in a, estimate the partial pressure of $\text{O}_2(\text{g})$ in the atmosphere at which iron pyrite (FeS_2) becomes thermodynamically unstable relative to iron II oxide (FeO). What does this imply about the content of oxygen in the early atmosphere of the Earth? (Note that there are other factors involved, including the rate of reaction 3.1 and the temperature dependence of the equilibrium constant, but the argument based on the thermodynamics of the reaction remains valid).

Substance	ΔH° (kJ/mol)	ΔG° (kJ/mol)	S° (J/mol·K)
FeO(s)	- 272.04	- 251.43	60.75
FeS ₂ (s)	- 167.36	- 156.16	53.86
O ₂ (g)	0.0	0.0	205.14
S(s)	0.0	0.0	31.80

5) Simple calculations can often tell us interesting information about planetary atmospheres.

a) Based on the information in the Chapter 1 handout, find the total mass of the atmosphere of Earth and Venus. Compare the result obtained for the Earth to the value given in the Chapter 1 notes. (HINT: The force exerted by a mass m in the gravitational field of a planet is $F = mG_s$, where m is mass and G_s is the gravitational constant for the planet. Force is also related to pressure by the relationship $p = F/A$, where A is area.)

b) Use your result from a and the information on atmospheric composition to find the total mass of argon in the atmosphere of the Earth and Venus. Comment on your results.

6) For an adiabatic expansion of an ideal gas, it may be shown that

$$(T_i/T_f)^\gamma = (p_i/p_f)^{\gamma-1} \quad (6.1)$$

where $\gamma = C_p/C_v$ is the ratio of the constant pressure heat capacity to the constant volume heat capacity of the gas. Eq 6.1 assumes that over the temperature range of the expansion γ is constant.

a) According to the barometric equation

$$p_z = p_0 \exp(-z/H) \quad (6.2)$$

where p_z is atmospheric pressure at altitude z and p_0 is atmospheric pressure at some reference altitude (taken to be sea level for the Earth).

By combining eq 6.1 and 6.2 find an expression for T_z , the temperature for an air parcel initially at sea level and temperature T_0 after it rises to an altitude z , and assuming that it expands adiabatically as it rises. Give your expression in terms of p_0 , T_0 , γ , H , z , and/or other constants. (HINT: Let T_0 and p_0 be the temperature and pressure of the air at sea level, where $z = 0$. Then T_f and p_f are equivalent to T_z and p_z , the temperature and pressure at altitude z .)

b) Using the following values for the terms in your expression for T_z , find the value for $T_z - T_0$ for the case $z = 1.0$ km. Note that this value for $T_z - T_0$ corresponds approximately to the adiabatic lapse rate for dry air in the Earth's atmosphere.

$$\gamma = 1.40 \quad T_0 = 288. \text{ K} \quad H = 7.4 \text{ km} \quad p_0 = 1.00 \text{ atm}$$